

Terra Incognita となる解像度のための Mellor-Yamada モデルの拡張

An extension of Mellor-Yamada model to apply for the resolution of
Terra Incognita

伊藤純至(東大AORI/気象研)

新野宏(東大AORI)

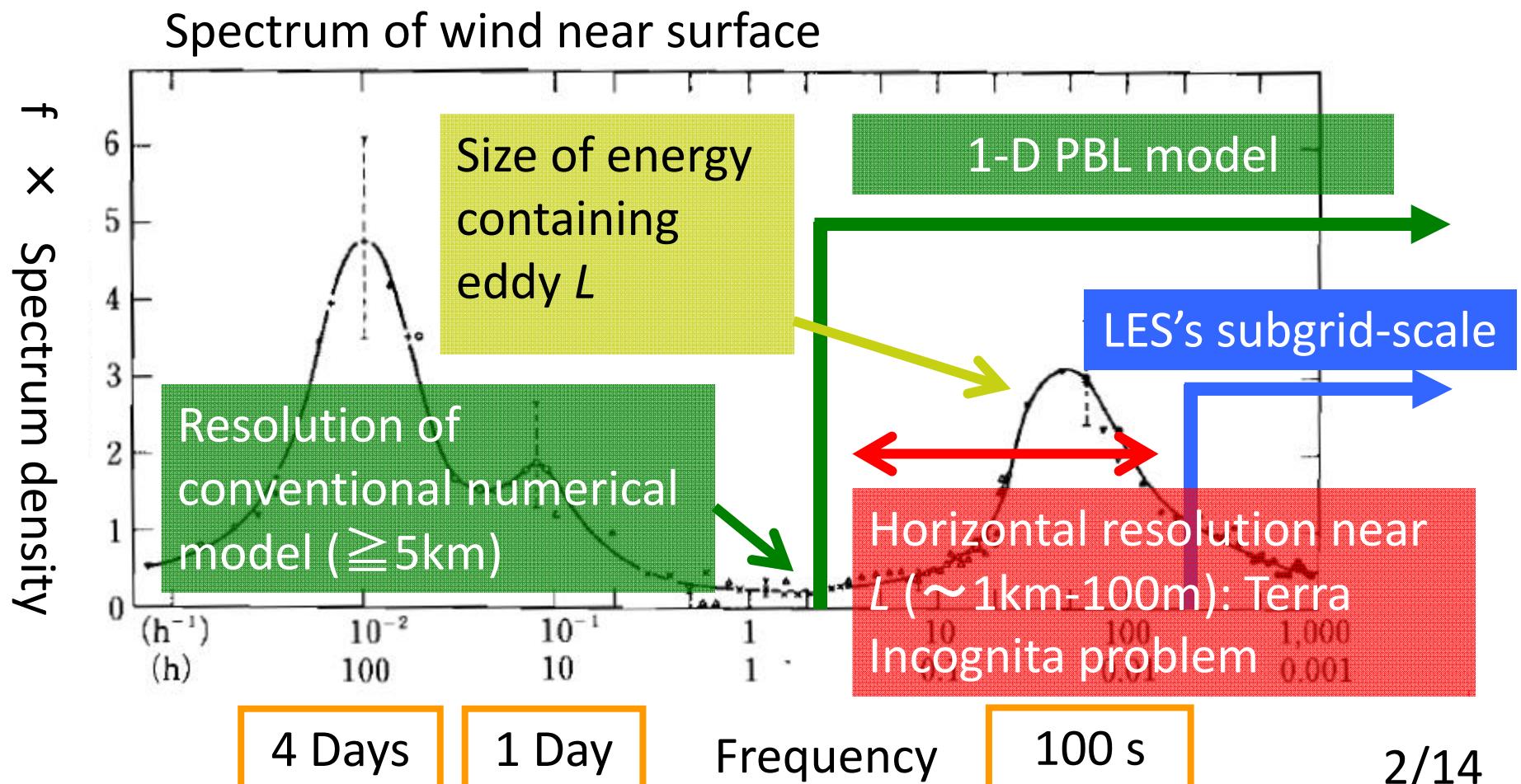
中西幹郎(防衛大)

Chin-Hoh Moeng(NCAR)

1. “Terra Incognita” (Grey Zone) problem in PBL turbulence modeling

※ Terra Incognita: 未知の領域

PBL: Planetary Boundary Layer 大気境界層



Objectives

Develop **a new PBL model** for resolution of Terra Incognita(TI)

- **Design based on LES reference**

e.g. Honnert et al. (2012); Shin and Hong (2013); Kitamura@MRI
(Submitted)

- **Test run** of a new model with horizontal resolution of TI

e.g. Ramachandran et al. (2011); Boutle et al. (2014)



An extension of UK model's non-local PBL scheme

- **Pragmatic extension** of Mellor-Yamada's PBL scheme used in JMANHM etc.

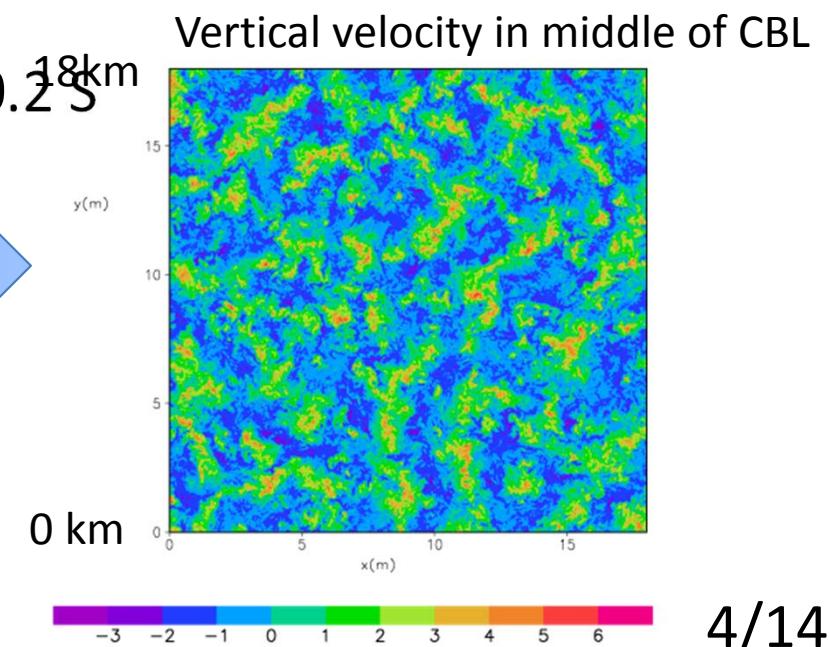
2. LES Reference

- Basic equation: Boussinesq approximation, Dry, Smagorinsky-Lilly's SGS
- Domain: $18\text{km} \times 18\text{km} \times 4.5\text{km}$
- Resolution: 25m
- Lateral Bound.: Doubly periodic
- Bottom Bound.: Bulk method (Momentum) / Heat flux
 $Q=0.2\text{K} \times \text{m/s}$
- Geostropic winds: $V=5 \text{ m/s}$
- $\Delta t:$



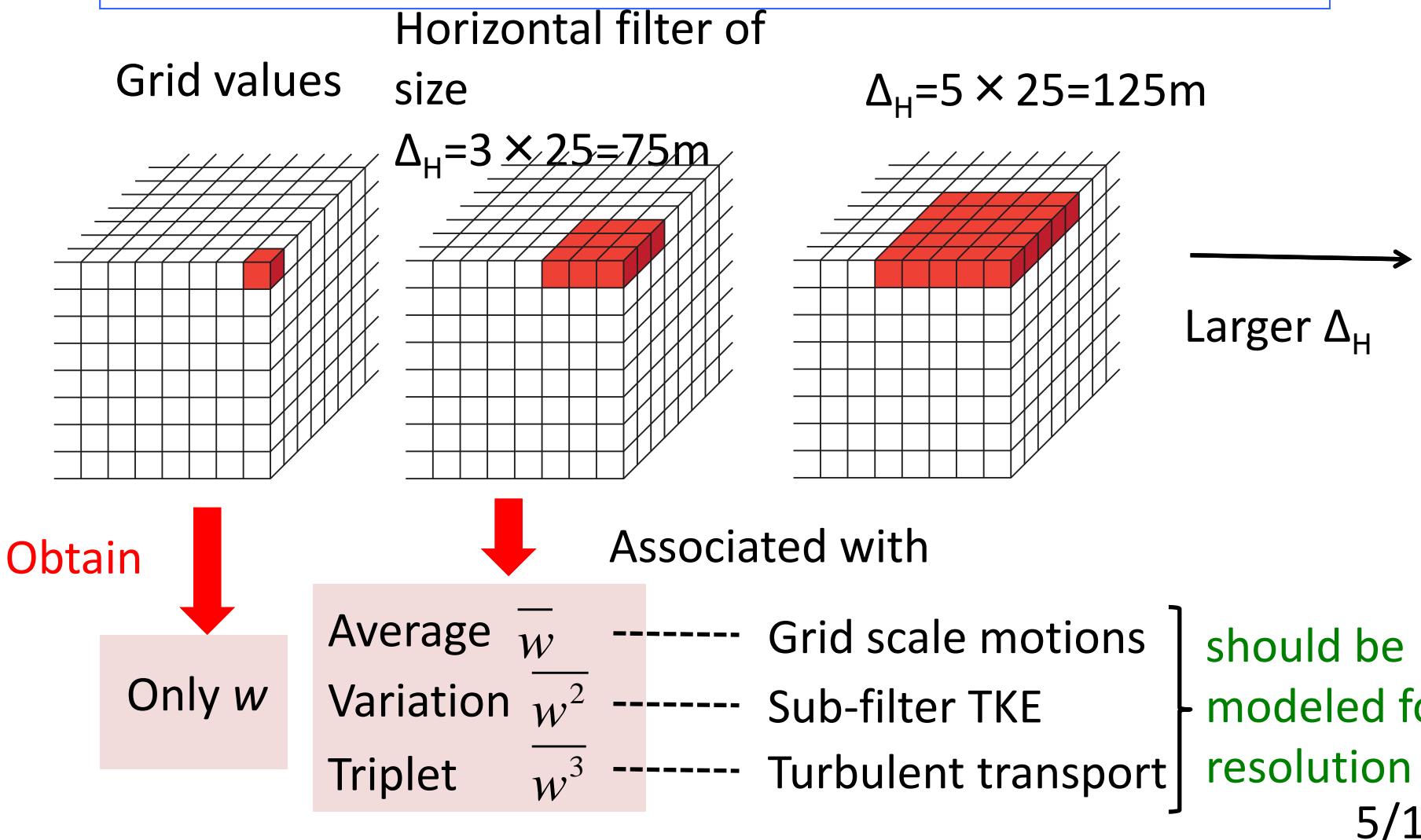
CBL (Convective Boundary Layer) grows

Sensitivity on Q and V (not shown)



Horizontal filter

Top-hat filter: $\overline{f(x_i)} = \iint f(x_i)G(x_i - x'_i)dx'_i$ $G(x_i - x'_i) = \begin{cases} 1/\delta_h & (|x_i - x'_i| \leq \delta_h) \\ 0 & (|x_i - x'_i| > \delta_h) \end{cases}$



Scaling and TKE budget

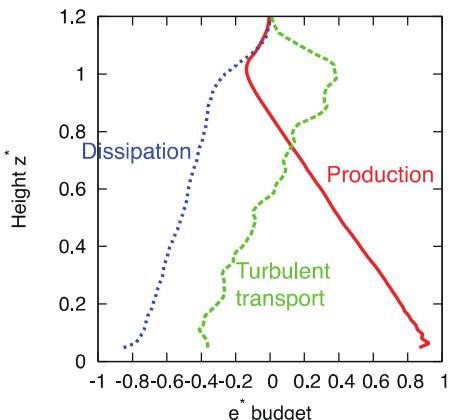
Free convection scaling (Deardroff 1970) is operated on results with “*”

- Length: height of convective boundary layer h
- Velocity : Convective velocity w_*
- Heat flux : Surface flux Q

Range $0.5 < \Delta_H^* (\equiv \Delta_H/h) < 2$ correspond to TI scale

MY model solves prognostic equation of TKE(乱流運動エネルギー):

$$\frac{\partial e}{\partial t} + u_i \frac{\partial e}{\partial x_i} = -\tau_{ij} S_{ij} + \frac{g}{\theta_0} f_3 - \frac{\partial}{\partial x_i} \left(\bar{u}_j e - \bar{u}_j \bar{e} \right) - \frac{1}{\rho_0} \left(\bar{u}_j p - \bar{u}_j \bar{p} \right) - \varepsilon$$

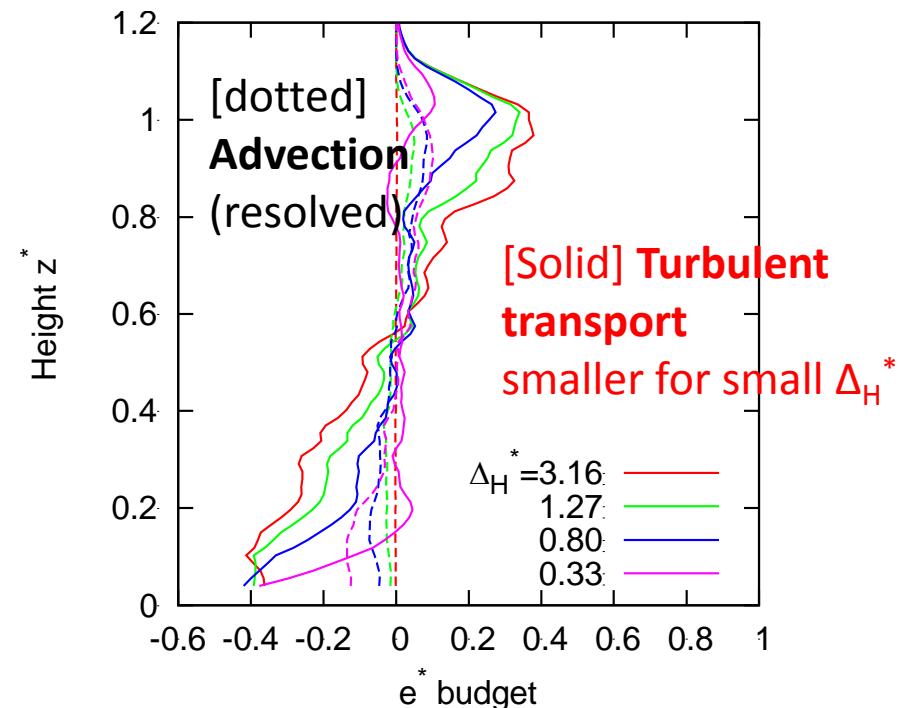
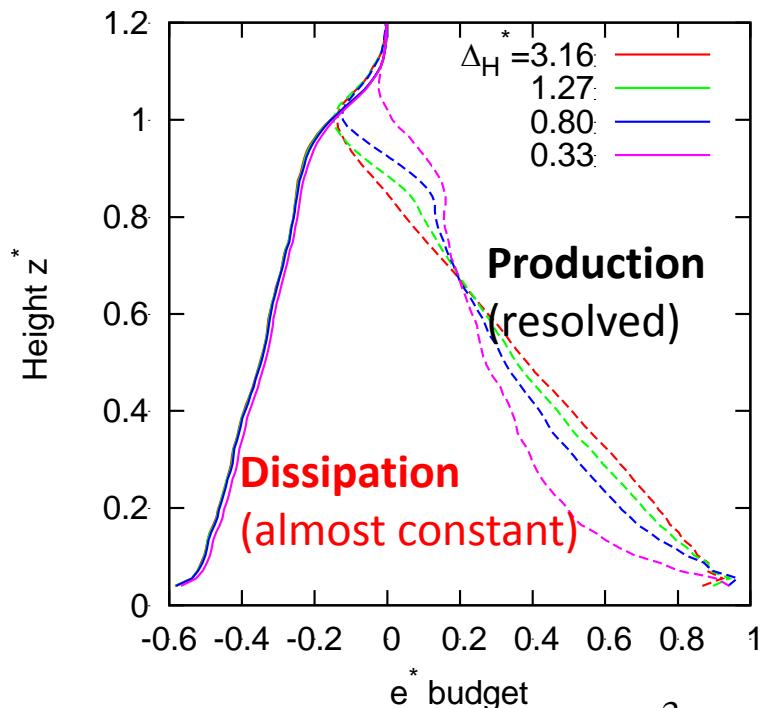


Vertical profiles of
over whole x-y cross
section

$\Delta_H^* \rightarrow \infty$: Meso-scale
limit

3. Design of model: Estimated terms in TKE budget for various Δ_H^*

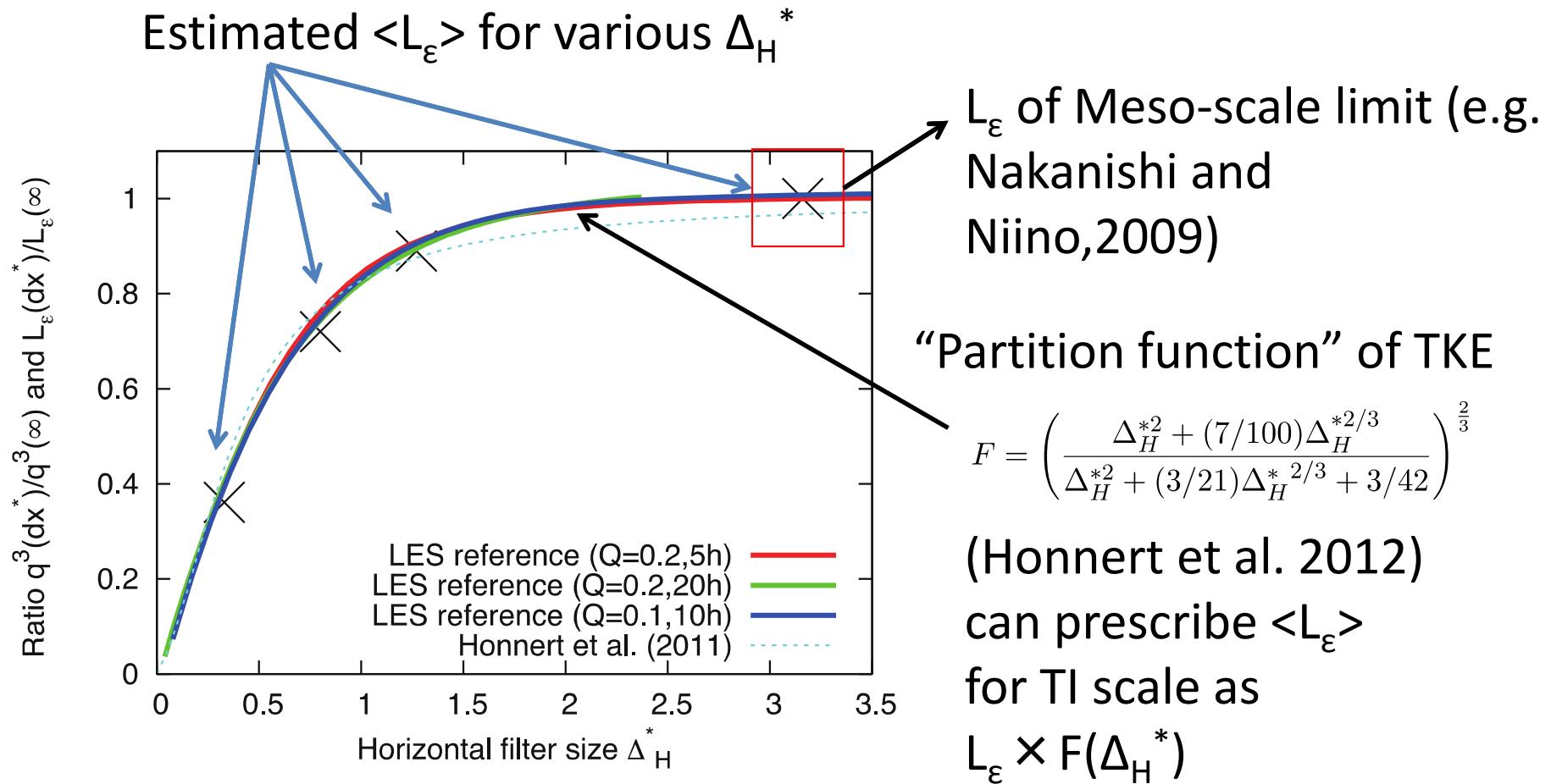
Vertical Profiles of terms in TKE budget (horizontal average)



① Dissipation: $\varepsilon = -\frac{q^3}{L_\varepsilon} \rightarrow q: \text{Turbulent velocity} \propto \text{TKE}$

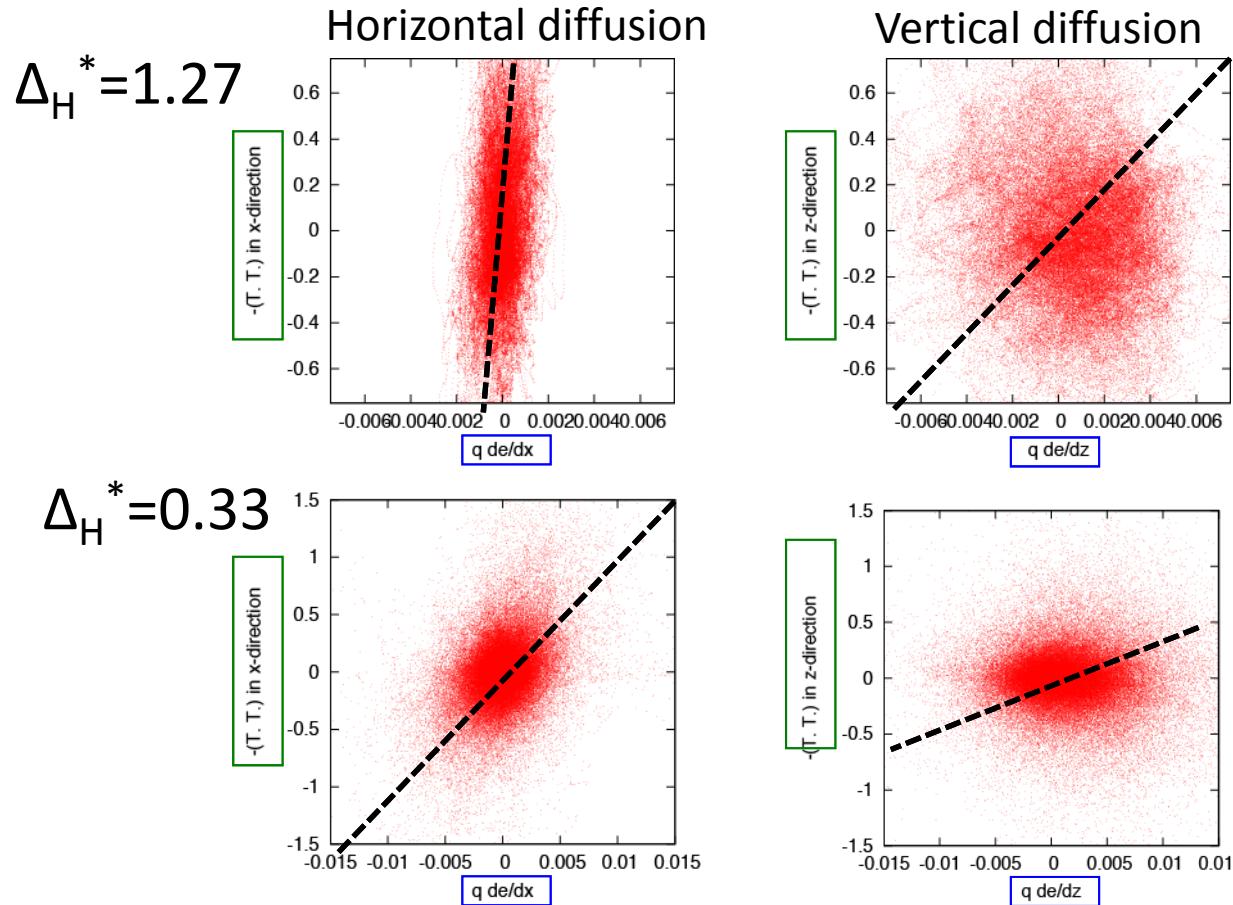
② Turbulent transport (T.T.): $T.T. = -q L_h \frac{\partial e}{\partial x_i}$ Assumed down-gradient form

① Horizontally averaged Dissipation length $\langle L_\varepsilon \rangle$



② Diffusion length

Turbulent transport = $-L_d q \partial TKE / \partial x_i$



Slope correspond to L_d

! Neglect counter-gradient diffusion plotted in 2nd and 4th quadrants

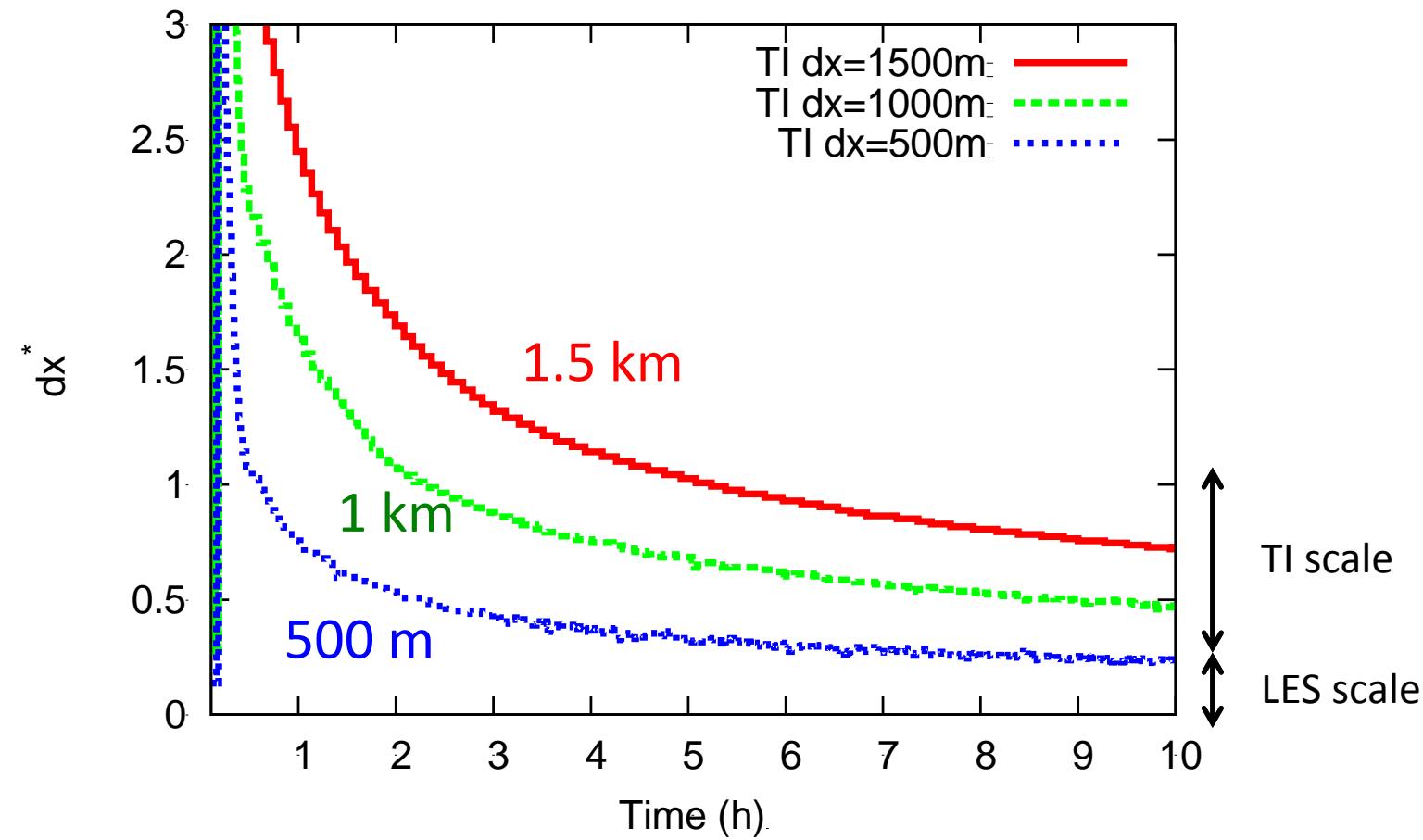
To reduce $L_d(\Delta_H^*)$ as $L_d(\infty) \times F(\Delta_H^*)$ seem reasonable

To rationalize the neglect of counter-gradient diffusion → A posteriori test

4. A posteriori test

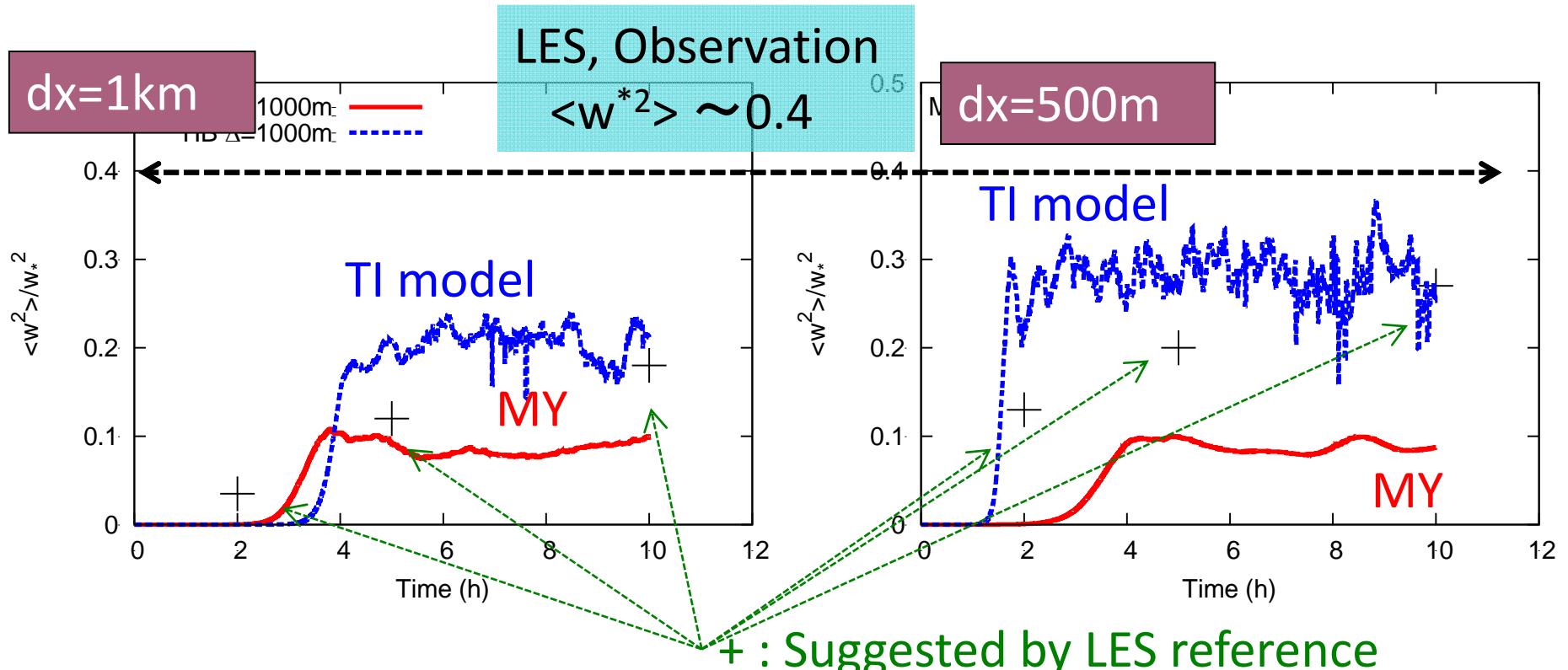
- Replace Smagorinsky's model in LES to
“TI model” | MY model (Nakanishi and Niino 2009)
- Grid number: $12 \times 12 \times 150$
- Horizontal resolution $dx=1.5\text{km}$ | 1km | 500m
Typical “TI scale” ↑
- Vertical resolution: $dz=25\text{m}$
- Environments are the same for LES reference to compare

Non-dimensional horizontal resolution dx^* ($\equiv dx/h$)



Strength of resolved convection w^{*2}

$\langle w^{*2} \rangle @ z=0.5h$



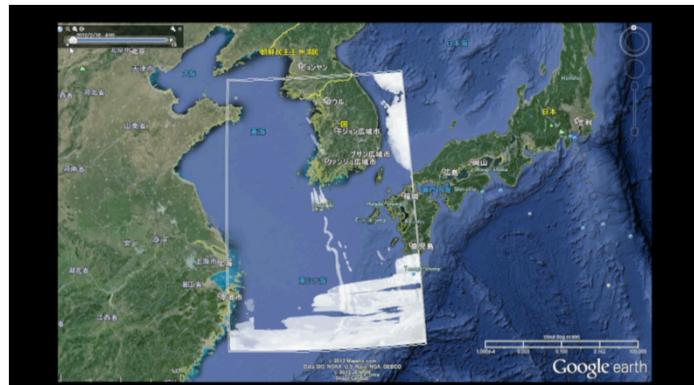
“TI Model” reproduce resolved convection in accord
with LES reference

Conclusion 1/2

- Design of “TI model” based on LES
 - **Almost completed** (To be submitted to BLM?)
- Implement in a numerical weather prediction (e.g. JMANHM, WRF) would be **very easy**
 - ↑ only reduce length scale in MY model
- We will **not** pursue further refinement (3rd order closure? stochastic forcing?)

Conclusion 2/2

- Current JMANHM with MY model **may behave qualitatively similar manner** to “TI model” (伊藤ら,2013年春気象学会), but it is unjustified
- Seeking verification by JMANHM with “TI model” ← Karman vortex shading was not good...

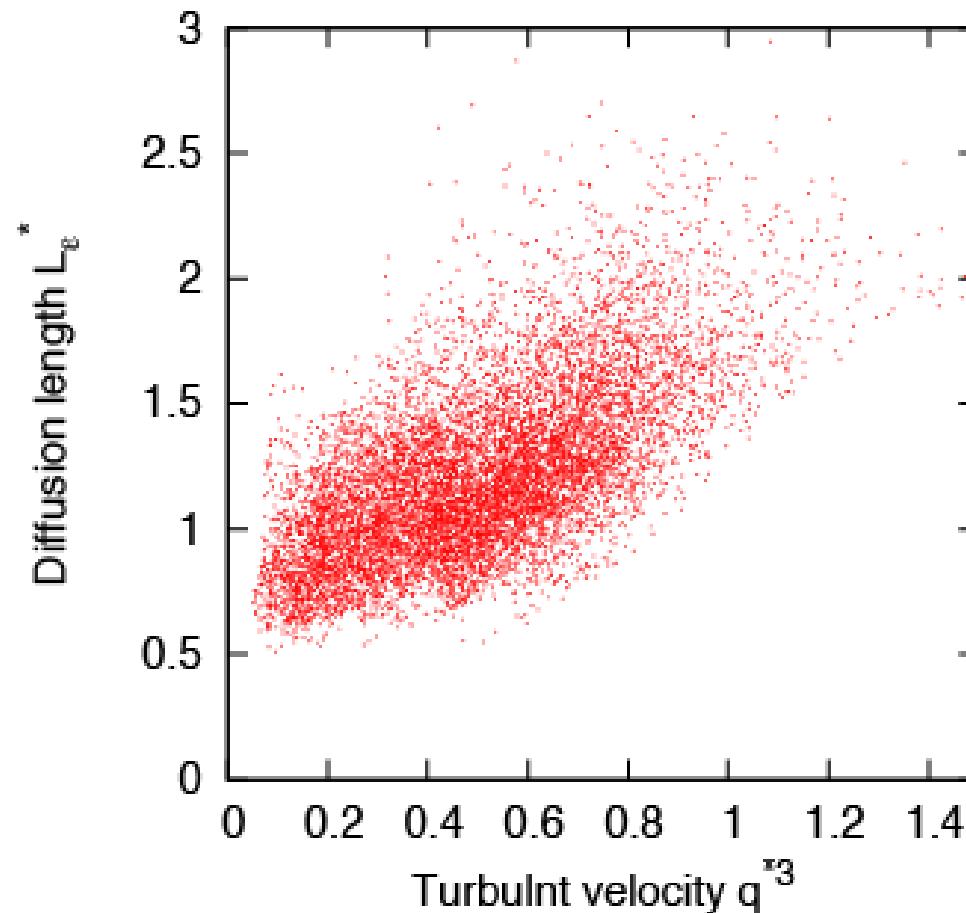


Simulated cloud

by JMANHM with a LFM configuration ($dx=2\text{km}$)

Local L_ε

L_ε without horizontal average @ $\Delta_H^*=0.8$, $z^*=0.5$



鉛直速度w@5h,z=0.5h

